The path to

QUANTUM GRAVITY

Francesca Vidotto



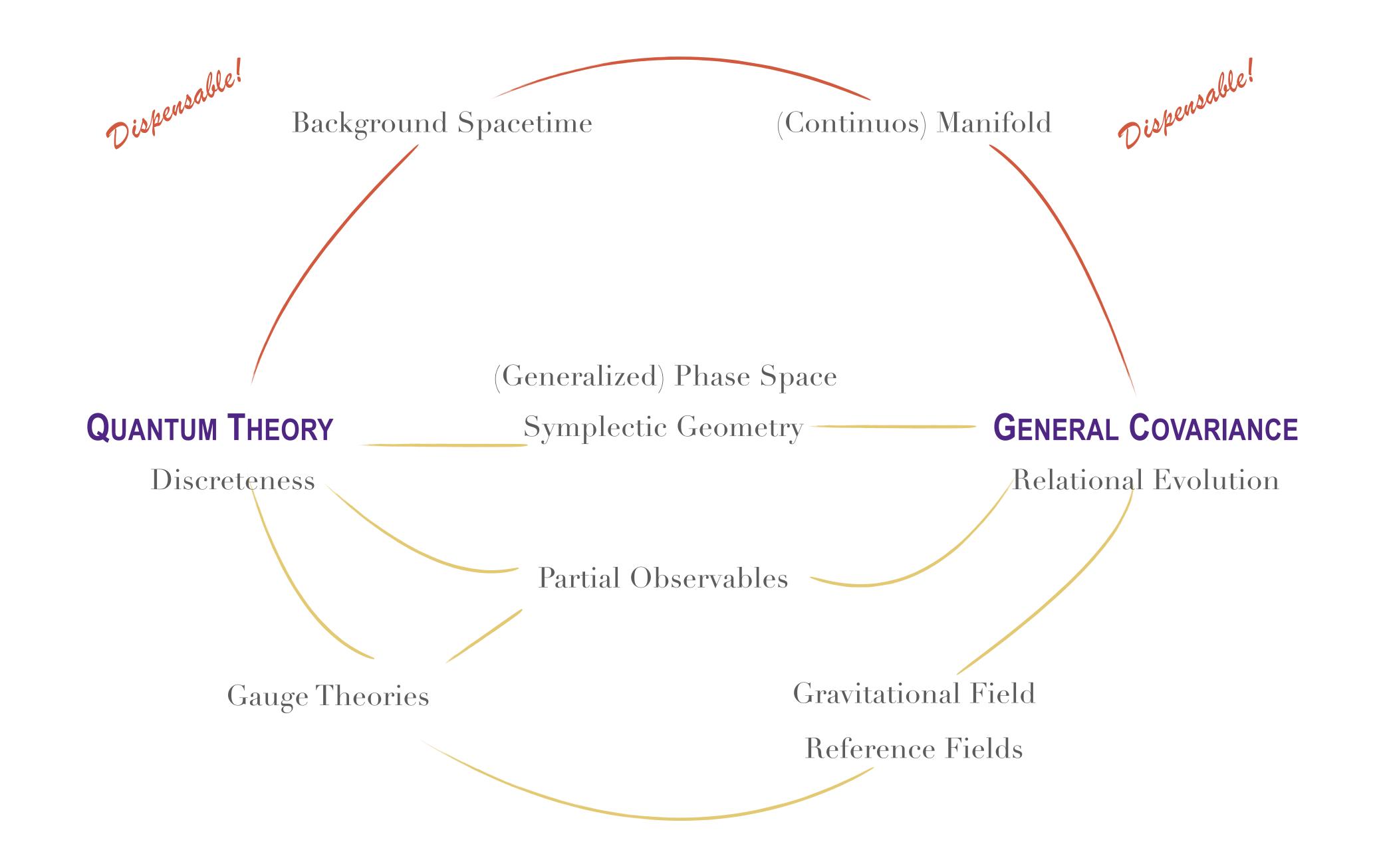




ENGAGING SCIENCE.

QUANTUM GRAVITY GOAL

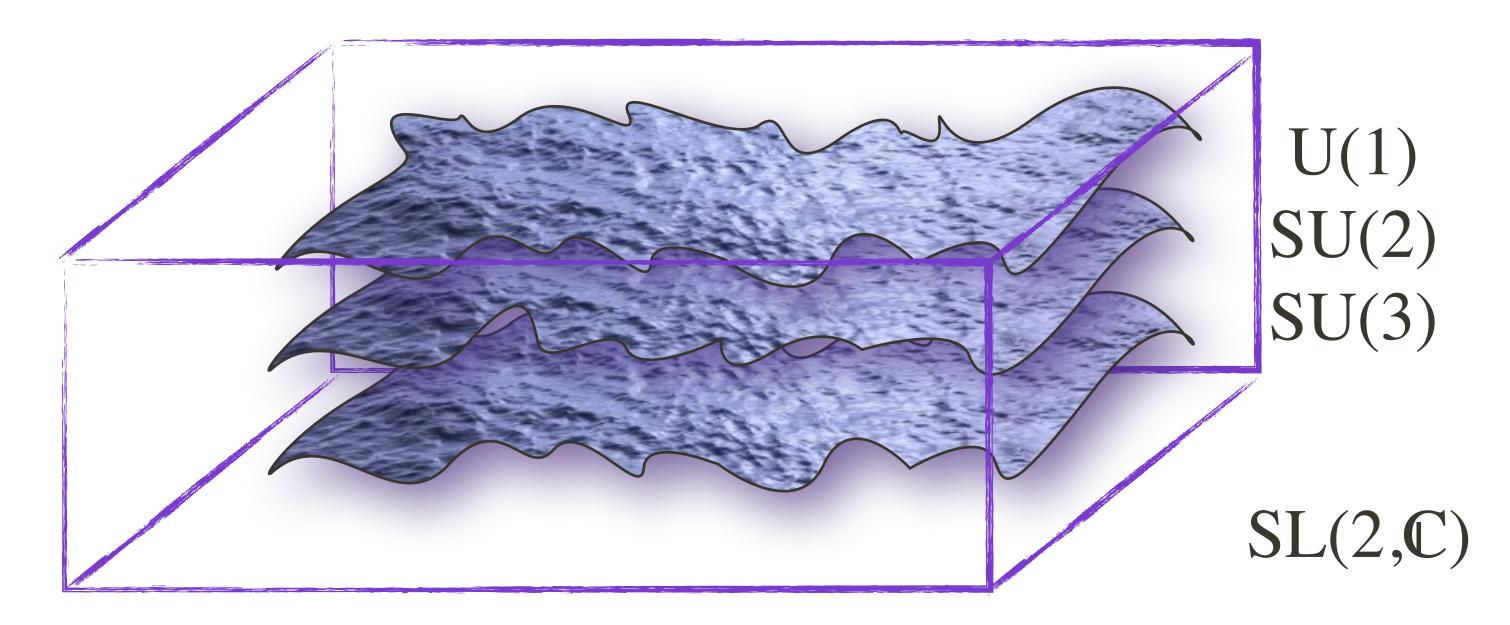
- A concrete realization of a quantum theory of the gravitational field ✓
- That is well defined without uncontrollable infinities •
- Whose classical limit is General Relativity
- In 4 Lorentzian dimensions
- With the standard matter couplings





EINSTEIN 1915

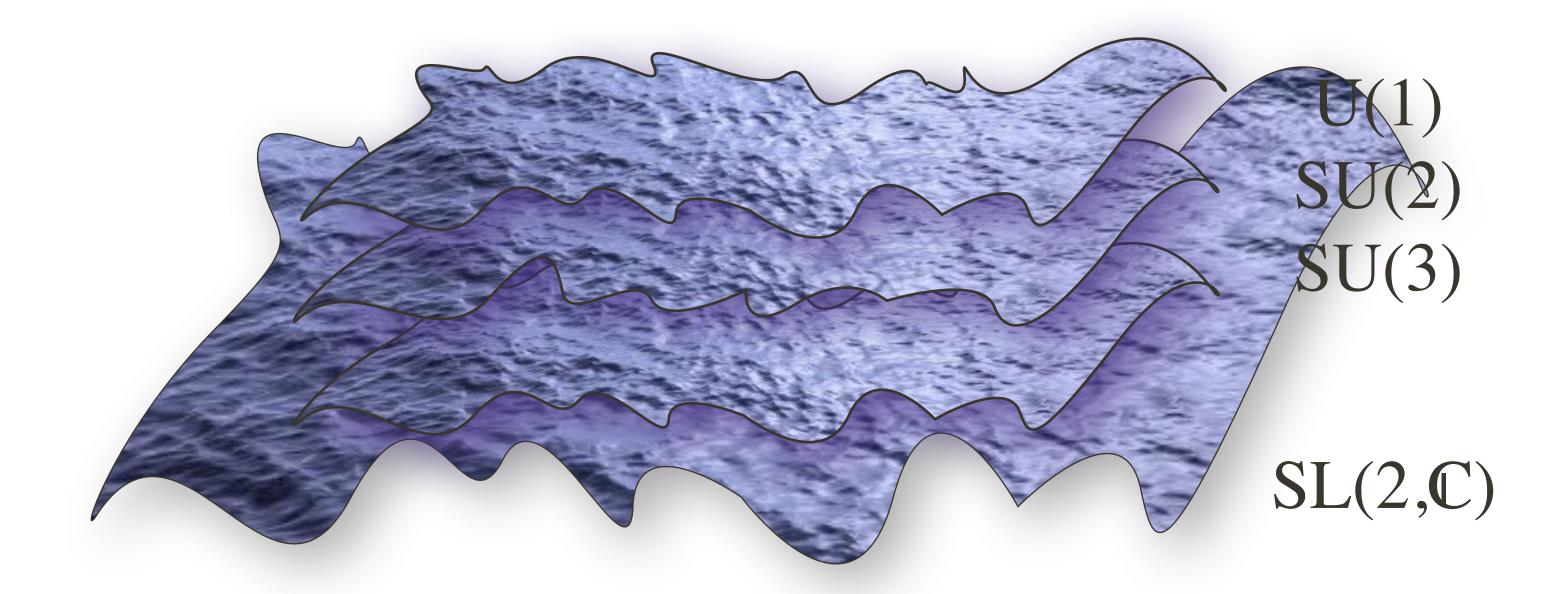
- ➤ General Relativity: spacetime is the gravitational field (substance)
- ➤ background independence / general covariance



Gravity as an interacting gauge field

EINSTEIN 1915

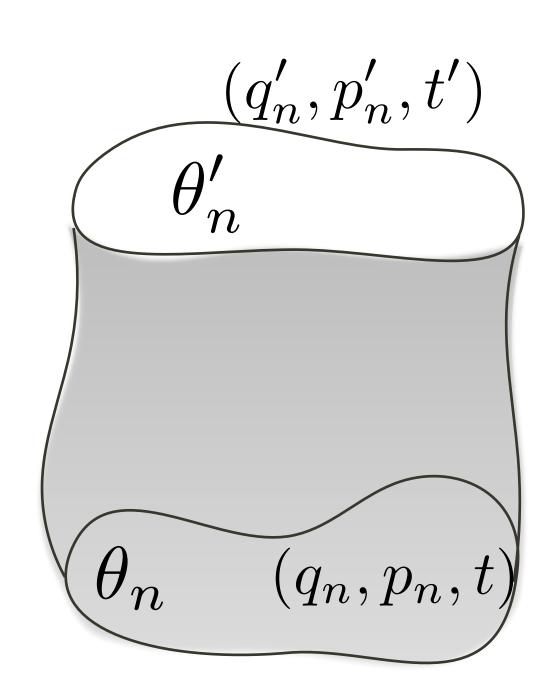
- ➤ General Relativity: spacetime is the gravitational field (substance)
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Gravity as an interacting gauge field

BOUNDARY FORMALISM

- ➤ Democratization of gauges: they all determine dynamical constraints among partial observables measured at the boundaries of a process
- ➤ The full content of a dynamical theory is in the constraints!
 - ➤ Yang-Mills constraint determines variable change wrt a change of the internal boundary frame.
 - ➤ Diff constraint determines variable change wrt a change in the location of the spatial boundary reference frame.
 - ➤ Hamiltonian constraint determines collective variable change wrt a change in the temporal location of the boundary (time of measurement)
- > Time is pure gauge, the Hamiltonian constraint determine time evolution
- ➤ Indeterminacy ←→ arbitrariness of the frame choice
- > Dynamics is the study of relations between partial observables that are gauge-dependent quantities of a system to which we can couple an apparatus



SPACETIME IS A PROCESS

QUANTUM MECHANICS

GENERAL RELATIVITY

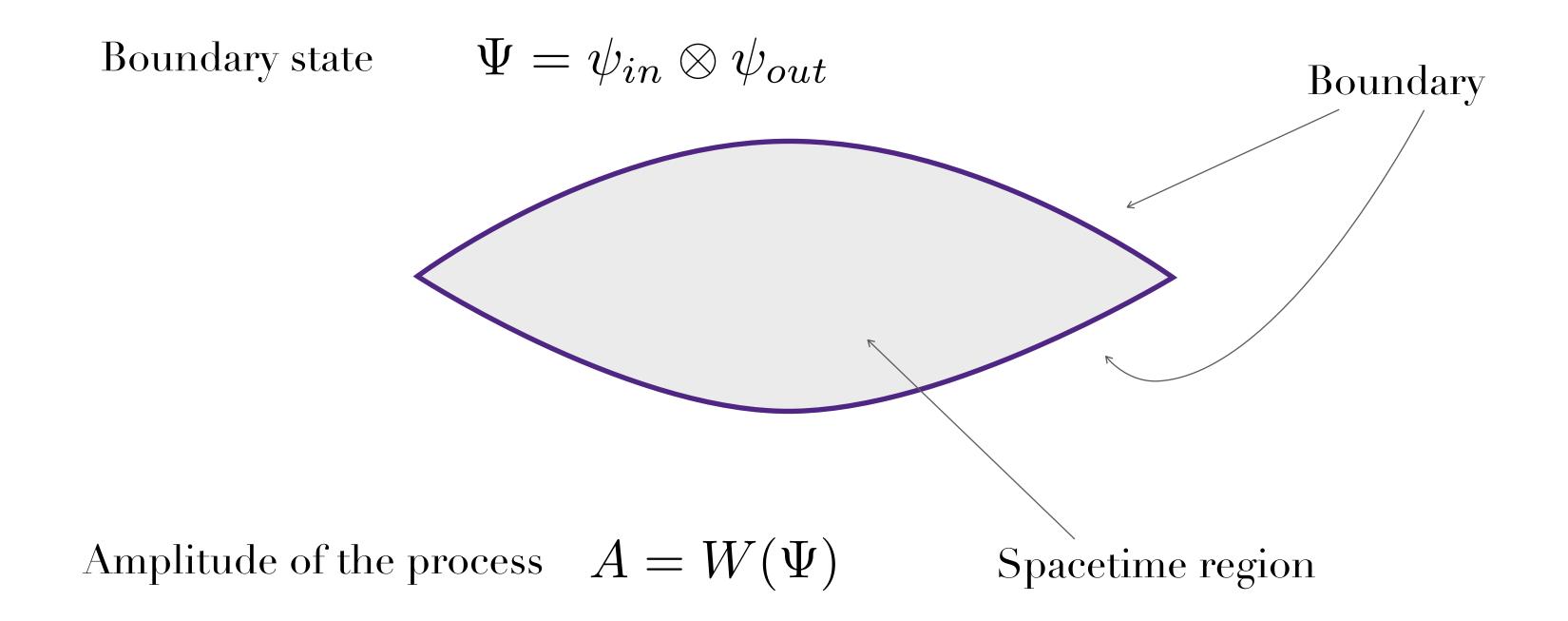
Process

State

← Locality →

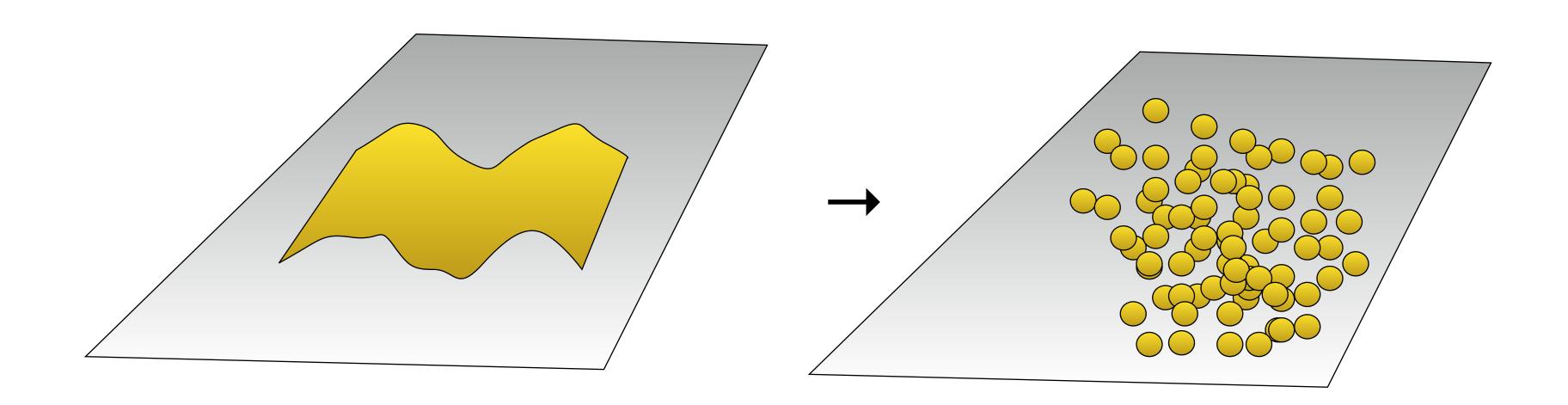
Spacetime region
Boundary, space region

Spacetime is a process, a state is what happens at its boundary.





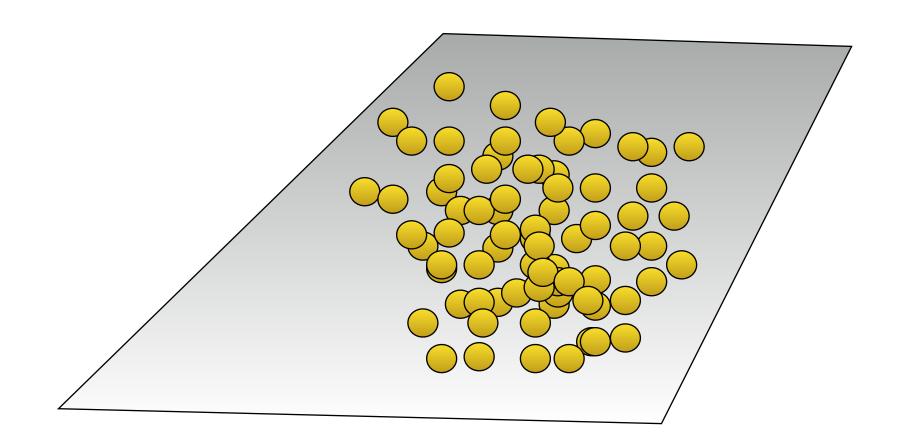
PRELIMINARY: RECALL Q.E.D.



The quanta of a field are particles (Dirac).

DISCRETENESS of the spectrum of the energy of each mode

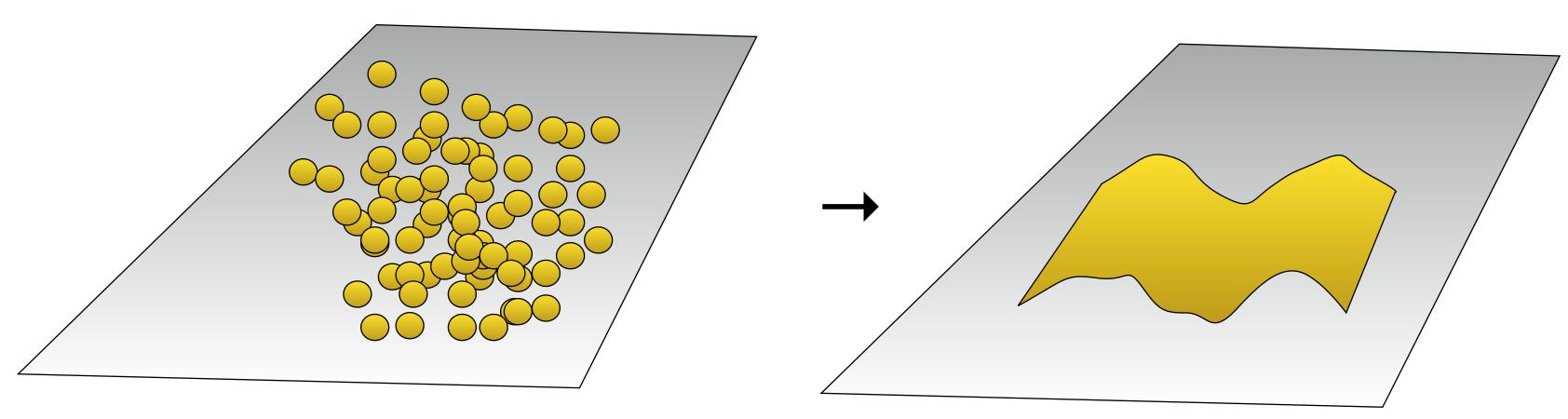
LIMIT $\hbar \longrightarrow 0$



- DISCRETE
- FUZZY
- PROBABILISTIC



LIMIT $\hbar \longrightarrow 0$



DISCRETE

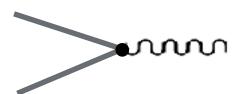
NO DISCRETENESS

FUZZY

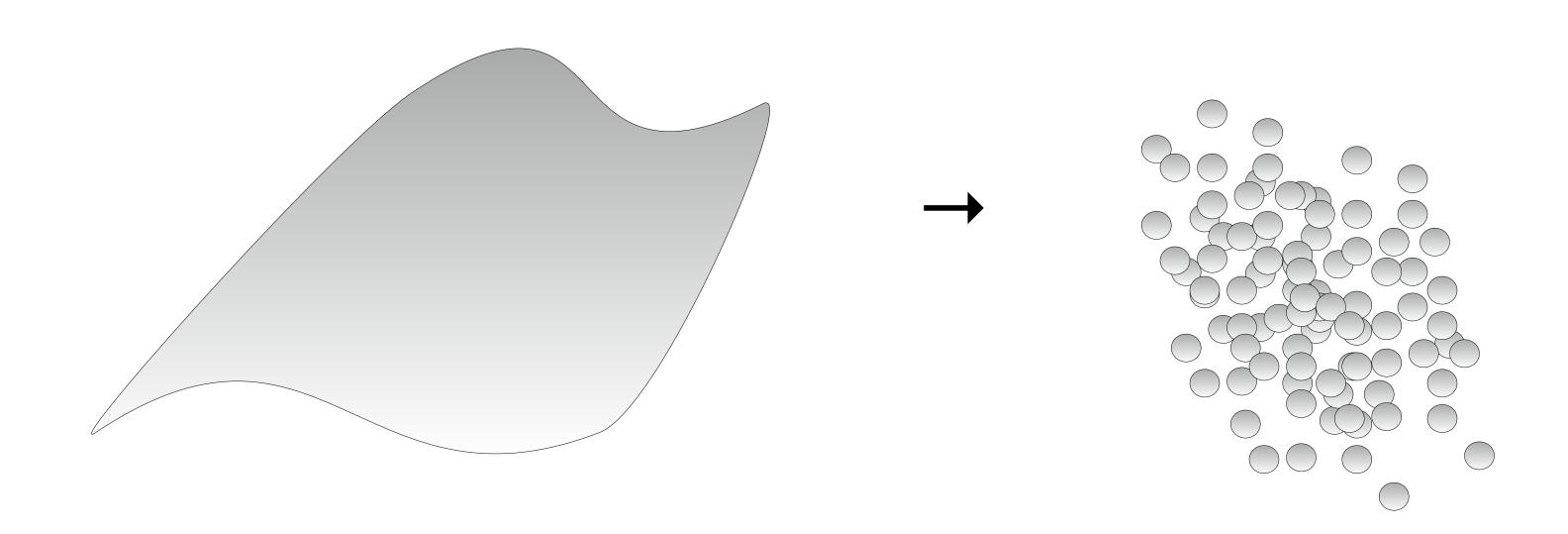
- NO FUZZYNESS
- $A_{\mu}(x) = \int d^3x \left(a(x)e^{+ikx}a^{\dagger}(x)e^{-ikx} \right)$

PROBABILISTIC

 \blacksquare A CLASSICAL FIELD $A_{\mu}(x)$



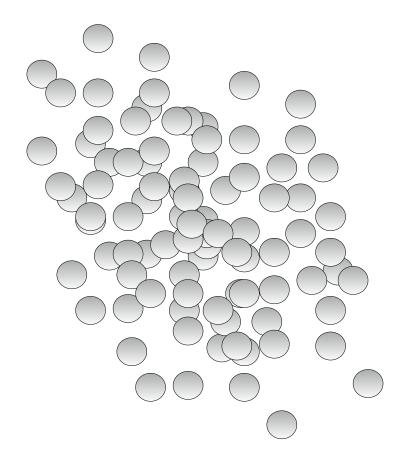
QUANTUM GRAVITY

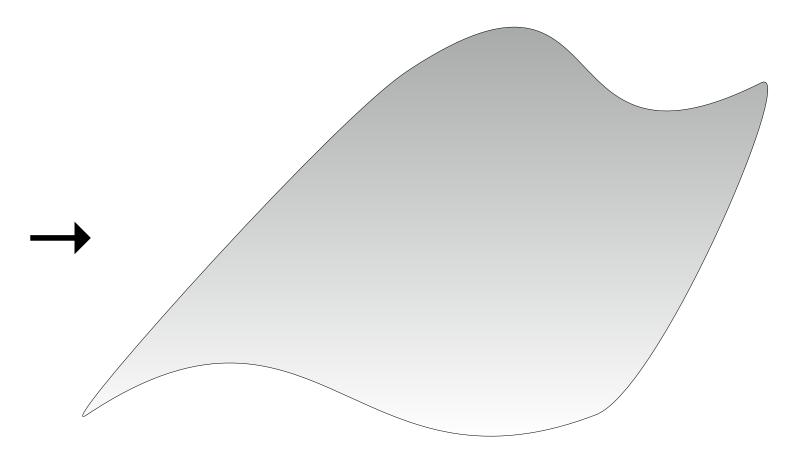


Quantum granularity of spacetime (Rovelli&Smolin 1995)

DISCRETENESS of the spectrum of geometrical operators such as volume

LIMIT $\hbar \longrightarrow 0$





- $\hspace{0.4cm} \hspace{0.4cm} \hspace{0.4cm}$
- FUZZY

NO FUZZYNESS

PROBABILISTIC

- A CLASSICAL FIELD
- $E_a^i(x) \to g_{\mu\nu}(x)$

FROM QUANTUM TO CLASSICAL

FROM QUANTUM TO CLASSICAL

The *classical* limit is $\hbar \longrightarrow 0$, the limit for ∞ quanta is relevant for the *continuous* limit. No thermodynamical limit is needed at that stage.

■ EMERGENCE OF SPACETIME IS STANDARD CLASSICAL EMERGENCE just as the electromagnetic field emerges from photons

SPACETIME IN THE QUANTUM REGIME IS MADE OF QUANTA

- there is no classical spacetime in the quantum regime
- same as in Q.E.D. where there are photons

SPACETIME IN THE QUANTUM REGIME IS A QUANTUM PROCESS

- states are defined by the continuity relations between quanta
- a spinfoam is a quantum interaction, but also a spacetime region

■ THERE IS NO TIME, THERE IS ONLY CHANGE in fact change is everything we measure!

QUANTUM GRAVITY IS THE DISCOVER OF A MINIMAL LENGHT

QUANTUM MECHANICS

Heisenberg Uncertainty

$$\Delta x > \hbar/\Delta p$$

Sharp localization requires large energy.

$$E \sim cp$$

GENERAL RELATIVITY

$$M \sim E/c^2$$

Black-Hole Horizon

$$R \sim GM/c^2$$

The horizon prevent a sharper localization.

$$\Delta x \geq R$$

QUANTUM GRAVITY

$$\ell_P = \sqrt{\frac{\hbar G}{c^3}} \sim 10^{-35} \, m$$

"Without a deep revision of classical notions it seems hardly possible to extend the quantum theory of gravity also to [the short-distance] domain."

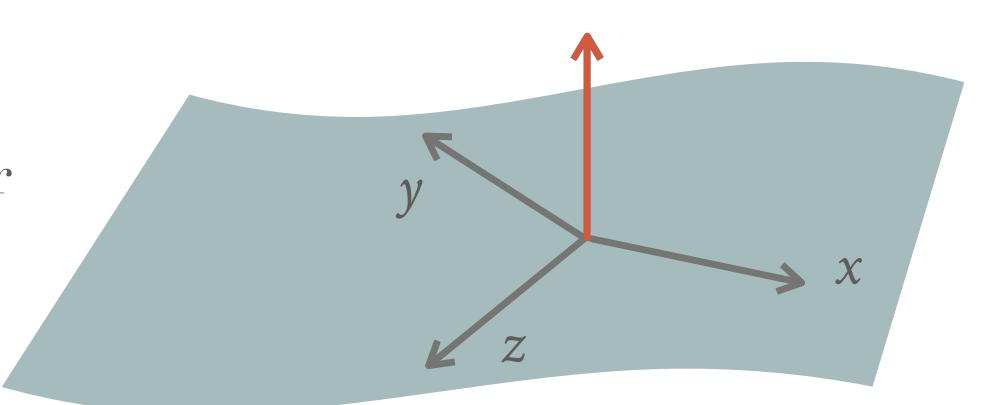
Matvei Bronstein



GENERAL RELATIVITY

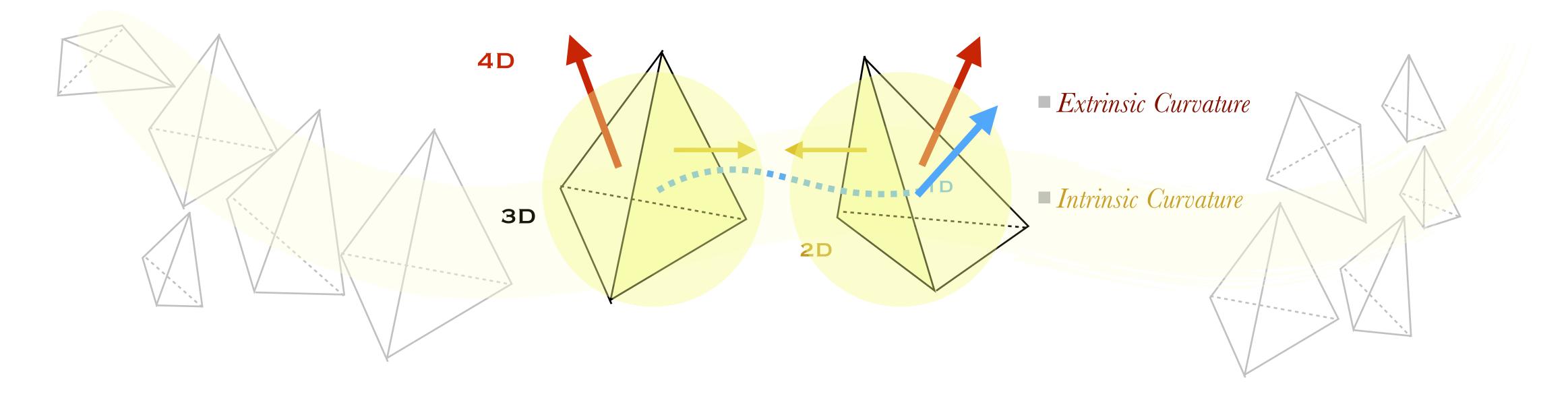
■ Tetrads (reference fields): required for coupling matter





- Hamiltonian formulation: (densitized) triads are conjugate to the Ashtekar connection
- Triads are rotation invariant: $so(3) \longrightarrow su(2)$

QUANTUM GEOMETRY



lacksquare Holonomy of the Ashtekar-Barbero connection along the link"

$$\vec{L}_{l} = \{L_{l}^{i}\}, i = 1, 2, 3 \text{ SU(2) generators}$$

$$gravitational field operator (tetrad)$$

$$Gauge invariant operator$$

$$G_{ll'} = \vec{L}_{l} \cdot \vec{L}_{l} \text{with} \qquad \sum G_{ll'} = 0$$

$$G_{ll'} = ec{L}_l \cdot ec{L}_{l\!\! ext{with}} \qquad \sum_{l\in n} G_{ll'} = 0$$

 A_l

Penrose's spin-geometry theorem (1971), and Minkowski theorem (1897)

REPRESENTING GEMETRIES

Composite operators:

Area:
$$A_{\Sigma}=\sum_{l\in\Sigma}\sqrt{L_l^iL_l^i}.$$
 Volume: $V_R=\sum_{n\in P}V_n, \quad V_n^2=rac{2}{9}\;|\epsilon_{ijk}L_l^iL_{l'}^jL_{l''}^k|.$

- Angle: $L_l^i L_{l'}^i$ $n \in \mathbb{R}$
- Geometry is quantized:
- eigenvalues are discrete
- the operators do not commute
- quantum superposition

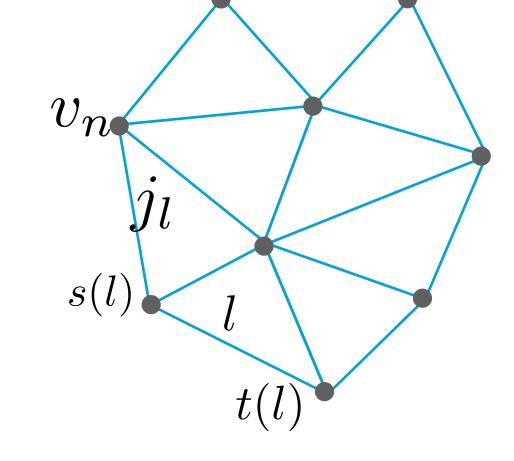
coherent states

Quantum states of space, rather than states on space.

GRAVITY AS A GAUGE THEORY

GAUGES IN GENERAL RELATIVITY
$$\longrightarrow$$
 $\left\{ \begin{array}{l} - \text{ diffeos} \\ - \text{ Lorentz} \end{array} \right\} \longrightarrow$ $\left\{ \begin{array}{l} - \text{ graph/lattice} \\ - \text{ group variables} \end{array} \right\} \longrightarrow$ LOOP QUANTUM GRAVITY

- ightharpoonup Abstract graphs: $\Gamma = N, L$
- \blacktriangleright Group variables: $\begin{cases} h_l \in SU(2) \\ \vec{L}_l \in su(2) \end{cases}$
- ➤ Graph Hilbert space: $\mathcal{H}_{\Gamma} = L_2[SU(2)^L/SU(2)^N]$
- \blacktriangleright The space \mathcal{H}_{Γ} admits a basis $|\Gamma, j_{\ell}, \nu_n \rangle$



➤ MANIFOLD is introduced as a device to connect with the classical limit

LOCALIZATION

> COUPLING TO A MATERIAL REFERENCE SYSTEM

The gravitational-field components with respect to the directions defined by the matter system are gauge-invariant quantities of the coupled system; but they are gauge-dependent quantities of the gravitational field, measured with respect to a given external frame.

> MEASUREMENT VS PREDICTIONS

Gauge variables are what we measure but we cannot predict.

A couple of gauge variables is gauge invariant and it allows predictions.

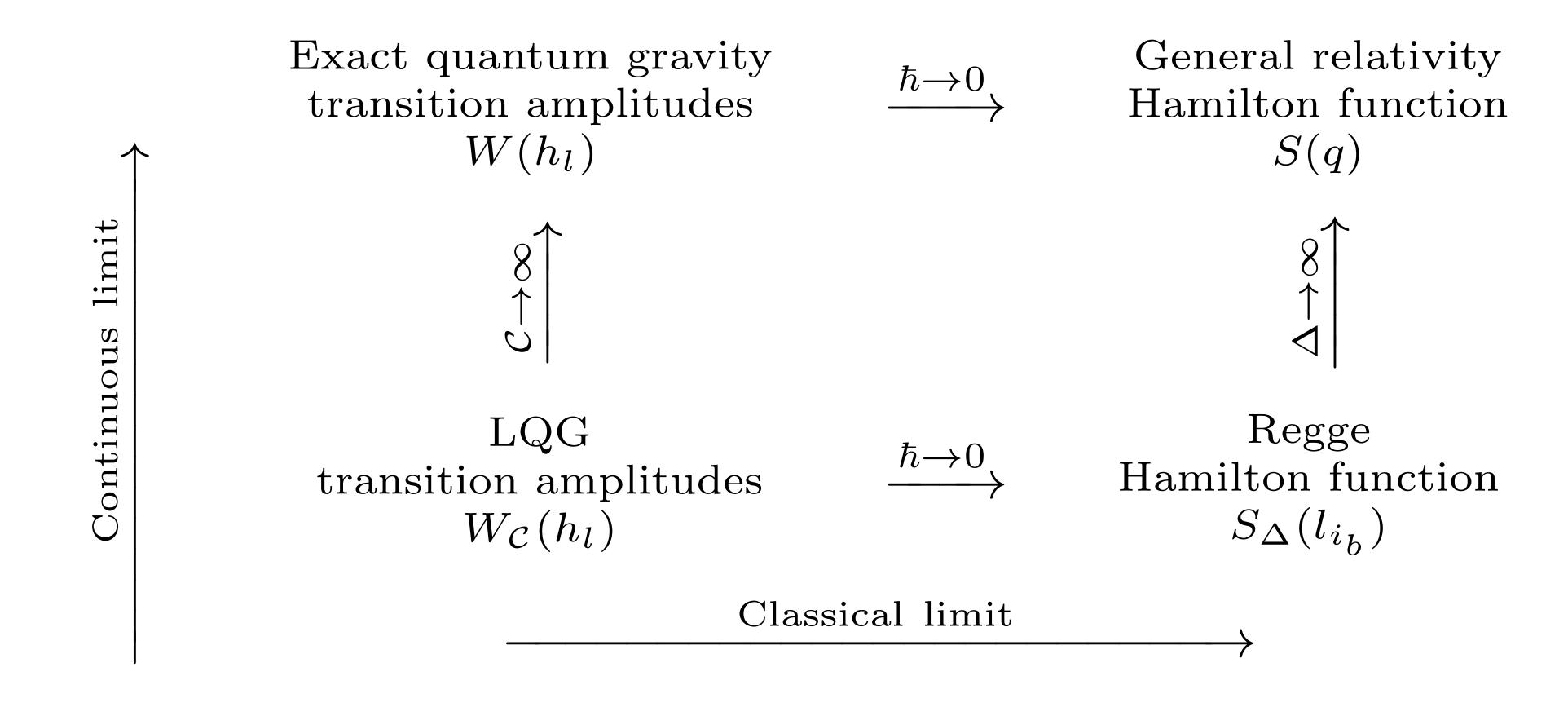
> RELATIONALITY

Bodies are only localized with respect to one another.

Bodies includes all dynamical objects, also the gravitational field.

Spacetime is built up by contiguity relations: being "next to one another".

STRUCTURE OF THE THEORY



No critical point

- QFT: critical phenomenon

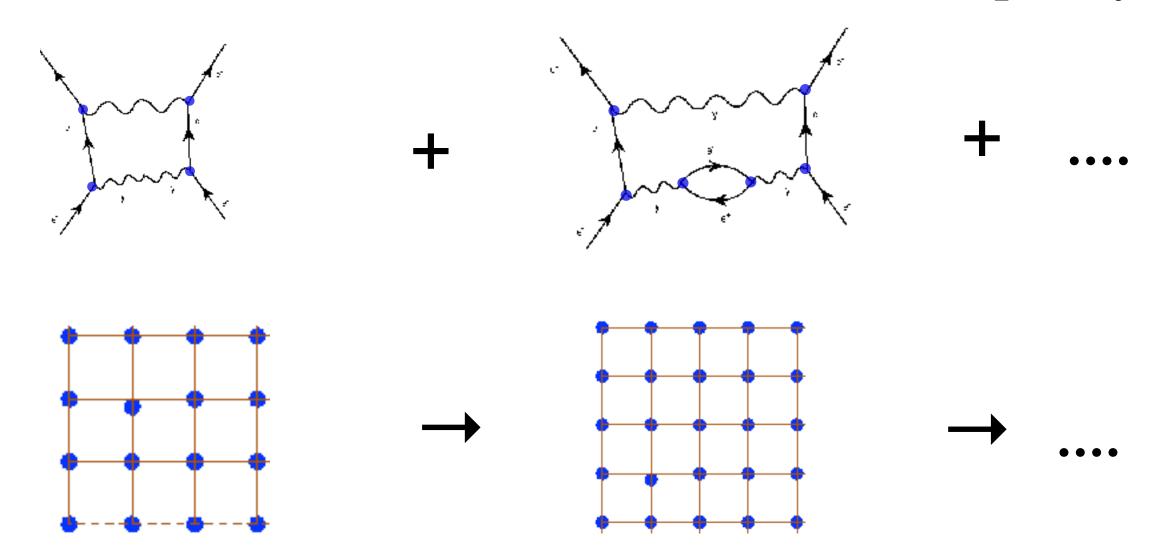
- Quantum Gravity: non-critical phenomenon

No infinite renormalization

Physical Scale ℓ_{Pl}

CONVERGENCE BETWEEN QED AND QCD

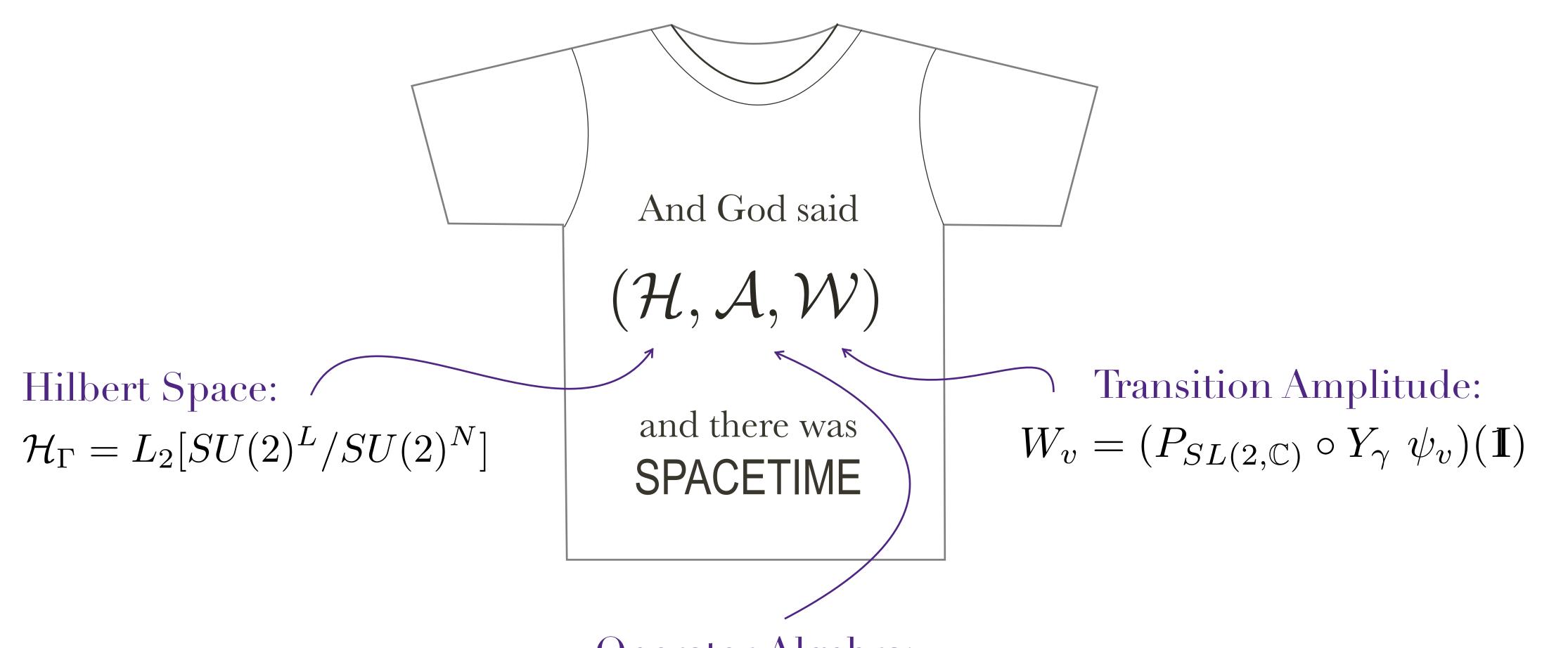
- ■All physical QFT are constructed via a truncation of the d.o.f. (QED: particles, QCD: lattice)
- All physical calculation are performed within a truncation.
- The limit in which all d.o.f. is then recovered is pretty different in QED qnd QCD:



Matter coupling as in LATTICE GAUGE THEORIES but without divergences!

- Quantum Gravity: Diff invariance!
- Lattice site = small region = excitations of the = quanta of space = quanta of the field

THE THEORY



Operator Algebra:

$$\left[L_a^i, L_b^j\right] = i\delta_{ab}\ell^2 \epsilon_k^{ij} L_a^k$$

CLASSICAL THEORY: TETRADS

Variables

$$e = e_a dx^a \in \mathbb{R}^{(1,3)}$$
 and $\omega = \omega_a dx^a \in sl(2,\mathbb{C})$

$$\omega = \omega_a dx^a \in sl(2, \mathbb{C})$$

Action

$$S[e,\omega] = \int B[e] \wedge F[\omega]$$
 where $B = (e \wedge e)^* + \frac{1}{\gamma}(e \wedge e)$

$$B = (e \wedge e)^* + \frac{1}{\gamma}(e \wedge e)$$

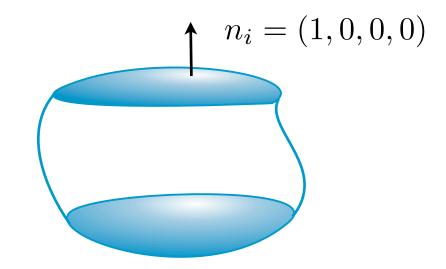
Boundary gauge s.t. tetrads are diagonal

$$B^{oi} = K^i = \frac{1}{\gamma} e^o \wedge e^i$$
 and $B^{ij} = L^i = e^o \wedge e^i$

$$B^{ij} = L^i = e^o \wedge e^o$$

Simplicity constraint

$$\vec{K} = \gamma \vec{L}$$

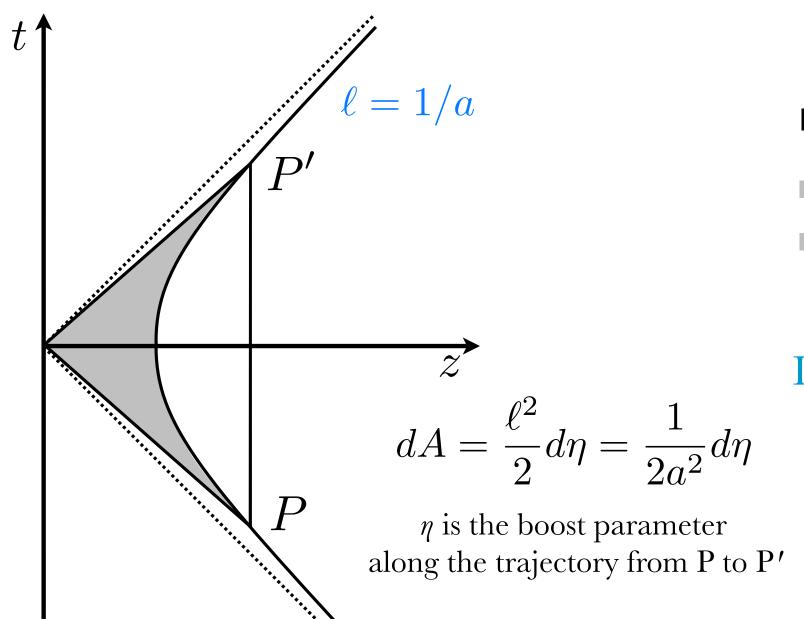


Lorentzian area

$$A = \int_{\mathcal{R}} e^{o} \wedge e^{i} = \int_{\mathcal{R}} \gamma K^{i} = \int_{\mathcal{R}} L^{i}$$

$$SL(2,\mathbb{C}) \to SU(2)$$

THERMODYNAMICS



- Constantly accelerated observer:
- K generator of boost
- $\blacksquare E = aK$ generator of proper time evolution

Linear simplicity constraint $B^{0k} = \gamma \epsilon^k{}_{ij} B^{ij}$

$$B^{0k} = \gamma \epsilon^k{}_{ij} B^{ij}$$

$$\vec{K} = \gamma \vec{L}$$

Simplicity Constrain Physics is defined locally!

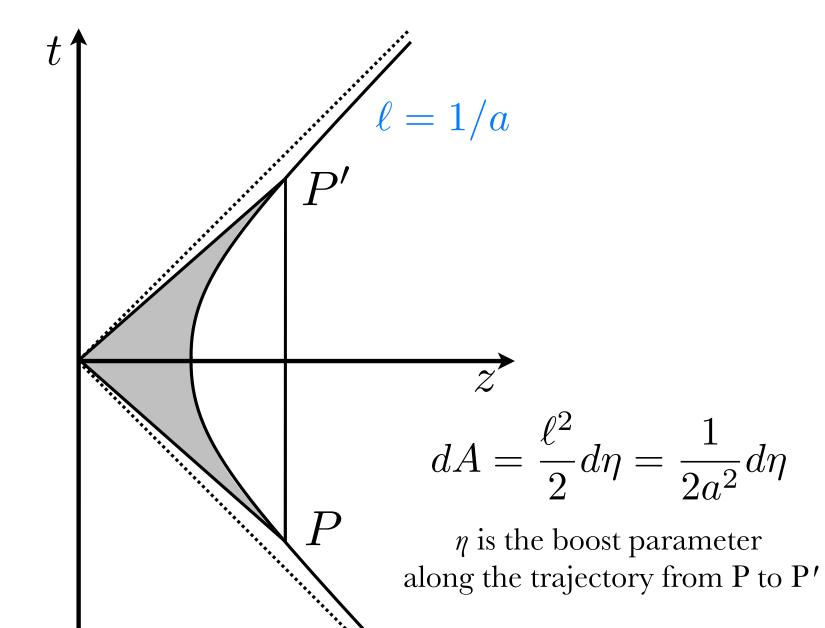
$$E = \frac{A}{8\pi G} l^{-1} \implies S_{\rm BH} = \frac{A}{4 G \hbar}$$

$$T = \frac{ha}{2\pi} \qquad H_A = 2\pi \sum_l K_l$$

$$\rho_A = e^{-H_A}$$

$$S_{ ext{EE}} = - ext{Tr}(
ho_A \log
ho_A) = 2\pi ext{Tr}(\sum_l K_l \;
ho_A)$$
 Bianchi 1204.5122 $S_{ ext{EE}} = rac{\mathcal{A}_\Sigma}{4G_0}$.

MAXIMAL ACCELERATION



- Constantly accelerated observer:
- K generator of boost
- $\blacksquare E = aK$ generator of proper time evolution

Linear simplicity constraint $B^{0k} = \gamma \epsilon^k{}_{ij} B^{ij}$

$$B^{0k} = \gamma \epsilon^k{}_{ij} B^{ij}$$

$$\vec{K} = \gamma \vec{L}$$

$$lacksquare$$
 Lorentzian area: $A = \int_{\mathcal{R}} \gamma K^z = \int_{\mathcal{R}} L^z$

$$A_{min} = 4\pi G\hbar$$

$$a_{max} = \sqrt{\frac{1}{8\pi G\hbar}}$$

$$\ell_{min} = \sqrt{8\pi G\hbar}$$

[Cainiello '81] [Cainiello, Gasperini, Scarpetta '91] [Bozza, Feoli, Lambiase, Papini, Scarpetta]

NO CURVATURE SINGULARITIES IN LQG



SPINFOAM AMPLITUDES

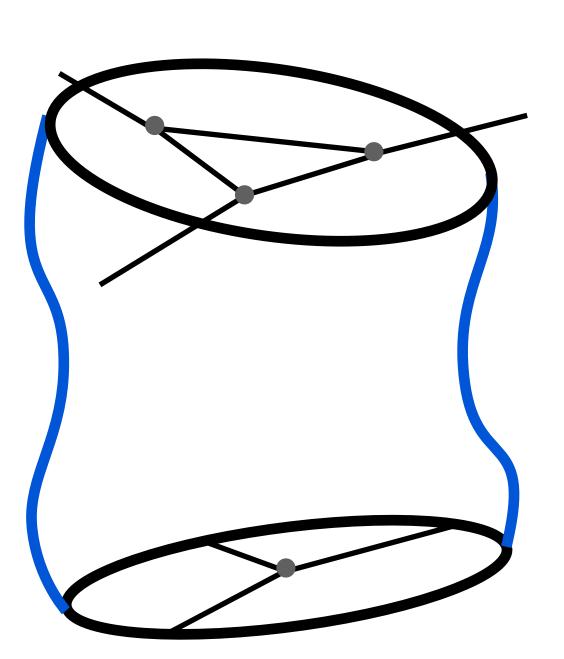
Probability amplitude $P(\psi) = |\langle W | \psi \rangle|^2$

for a state Ψ associated to the boundary of a 4d region

$$W(q'_{ij}, q_{ij}) \sim \int_{\partial g = q', q} Dq e^{iS}$$

- Superposition principle $\langle W|\psi\rangle=\sum_{\sigma}\ W(\sigma)$
- Locality: vertex amplitude $W(\sigma) \sim \prod_{v} W_v$.
- Lorentz covariance $W_v = (P_{SL(2,\mathbb{C})} \circ Y_\gamma \ \psi_v)(\mathbb{I})$
- UV and IR finite (with Λ)
- Classical limit: GR (with Λ) (via Regge discretization)

 Barrett et al. '09



The path to

QUANTUM COSMOLOGY

Francesca Vidotto







ENGAGING SCIENCE.

IN THIS LECTURE

■ THEORY: Loop Quantum Gravity

Covariant Loop Quantum Gravity, with C. Rovelli

■ SETTING: Primordial Universe

Relational Quantum Cosmology

■ APPLICATION: Quantum State of the Primordial Universe

Primordial fluctuations from quantum gravity, with F. Gozzini

Quantum Regime

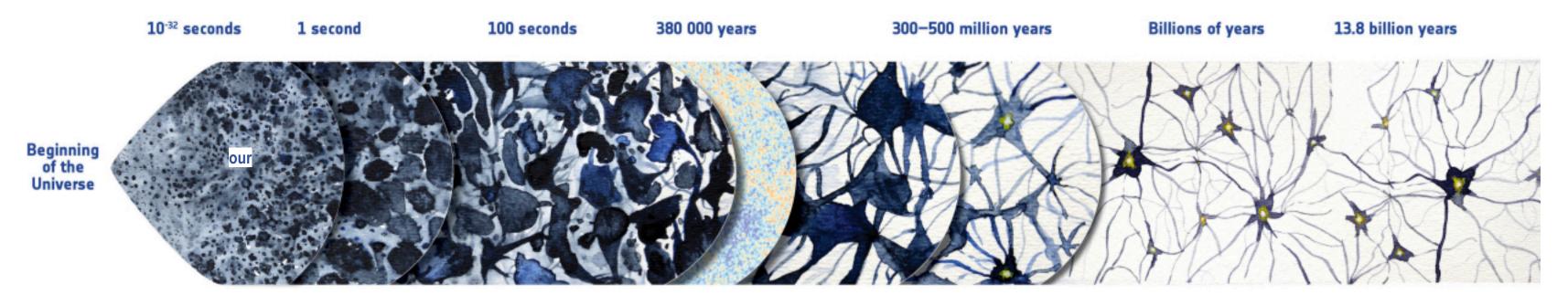


Image credit: ESA

■ Graph Approximation

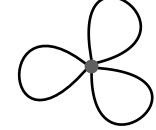
■ Hartle-Hawking State

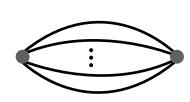
Transition Amplitude

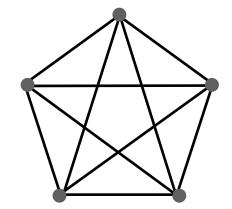
■ Computational Method

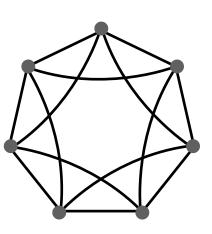
Quantum Regime











Hartle-Hawking State

■ Transition Amplitude

Computational Method

Quantum Regime

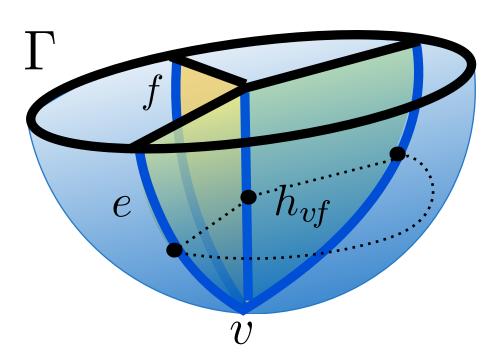
■ Graph Approximation

■ Hartle-Hawking State

Transition Amplitude

Computational Method

Spinfoam Hartle-Hawking state



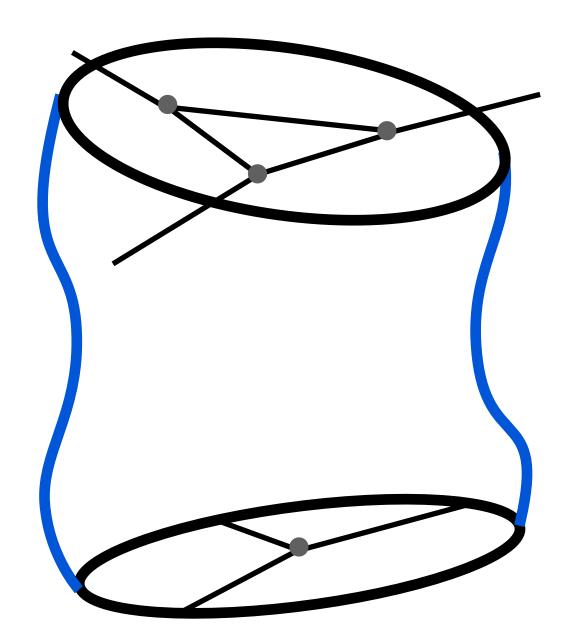
Quantum Regime

■ Graph Approximation

Hartle-Hawking State

Transition Amplitude

Computational Method



Quantum Regime

■ Graph Approximation

■ Hartle-Hawking State

Transition Amplitude

■ Computational Method

$$W_{\mathcal{C}}(h_l) = \int_{SU(2)} dh_{vf} \prod_{f} \delta(h_f) \prod_{v} A(h_{vf})$$

1ST-ORDER FACTORIZATION

Classical dynamics
$$H = const (a\dot{a}^2 - \frac{\Lambda}{3}a^3) = 0$$
 $\dot{a} = \pm \sqrt{\frac{\Lambda}{3}}a$

1ST-ORDER FACTORIZATION

■ classical dynamics
$$S_H = const \int dt \left(a\dot{a}^2 + \frac{\Lambda}{3}a^3\right)\Big|_{\dot{a} = \pm\sqrt{\frac{\Lambda}{3}}a} = const \frac{2}{3}\sqrt{\frac{\Lambda}{3}}(a_f^3 - a_i^3)$$

■ loop dynamics
$$\langle W|\psi_{H_{(z,z')}}\rangle = W(z,z') = W(z)\overline{W(z')}$$

1ST-ORDER FACTORIZATION

■ classical dynamics
$$S_H = const \int dt \left(a\dot{a}^2 + \frac{\Lambda}{3}a^3\right)\Big|_{\dot{a} = \pm\sqrt{\frac{\Lambda}{3}}a} = const \frac{2}{3}\sqrt{\frac{\Lambda}{3}}(a_f^3 - a_i^3)$$

loop dynamics
$$\langle W|\psi_{H_{(z,z')}}\rangle = W(z,z') = W(z)\overline{W(z')}$$

order (0)
$$= W_0(h_\ell, h_{\ell'}) = \delta_{\Gamma_\ell}(h_\ell, h_{\ell'})$$

ST-ORDER FACTORIZATION

■ classical dynamics $S_H = const \int dt \left(a\dot{a}^2 + \frac{\Lambda}{3}a^3\right)\Big|_{\dot{a} = \pm \sqrt{\frac{\Lambda}{3}a}} = const \frac{2}{3}\sqrt{\frac{\Lambda}{3}}(a_f^3 - a_i^3)$

loop dynamics

$$\langle W|\psi_{H_{(z,z')}}\rangle = W(z,z') = W(z)\overline{W(z')}$$

$$W_{\mathcal{C}_\infty}(z',z) = \int h_\ell \int h'_\ell \ \overline{\psi_{z'}(h'_\ell)} \ W_1(h'_\ell,h_\ell) \ \psi_z(h'_\ell)$$

order (1)
$$W_{\mathcal{C}_{\infty}}(z',z) = \int h_{\ell} \int h'_{\ell} \ \overline{\psi_{z'}(h'_{\ell})} \ W_{1}(h'_{\ell},h_{\ell}) \ \psi_{z}(h'_{\ell})$$

$$W_{1}(h'_{\ell},h_{\ell}) = \int_{SL(2,\mathbb{C})} \ \prod_{n=1}^{N-1} dG_{n} \ \prod_{\ell=1}^{L} \ P(h_{\ell},G_{\ell})P(h'_{\ell},G'_{\ell})$$

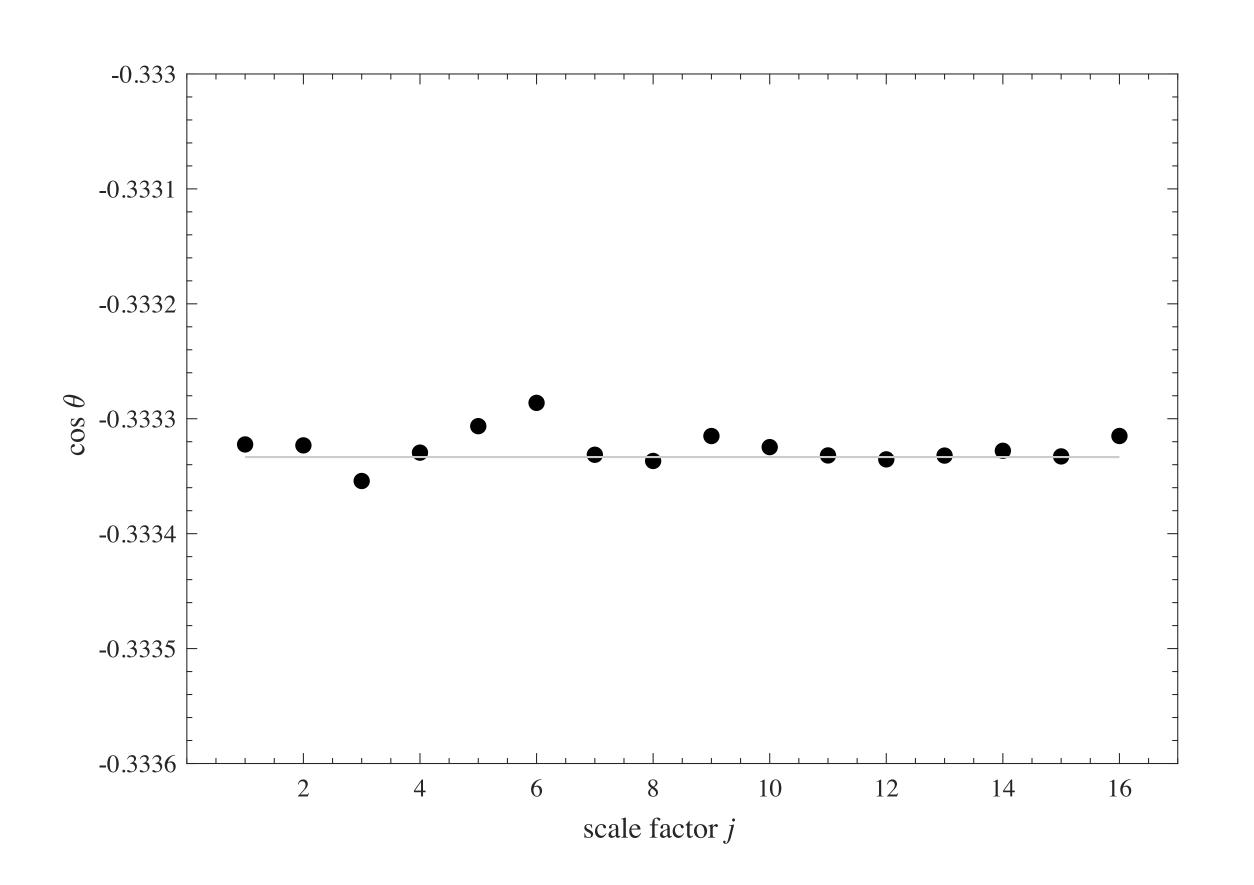
$$G_{\ell} = G_{n_s} G_{n_t}^{-1}$$

RECENT RESULTS

1. 3-sphere as emerging geometry

2. large fluctuations

3. large correlations

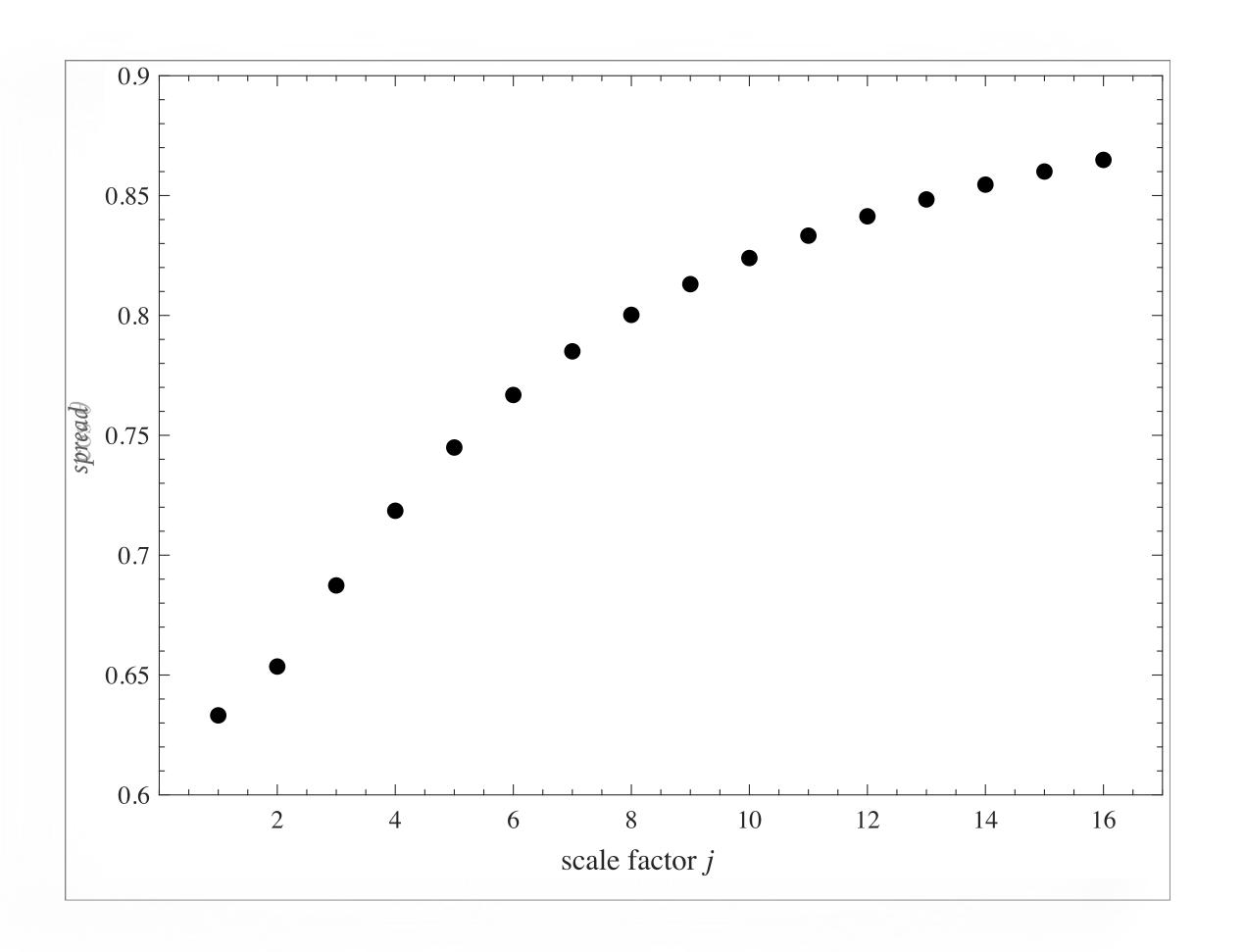


RECENT RESULTS

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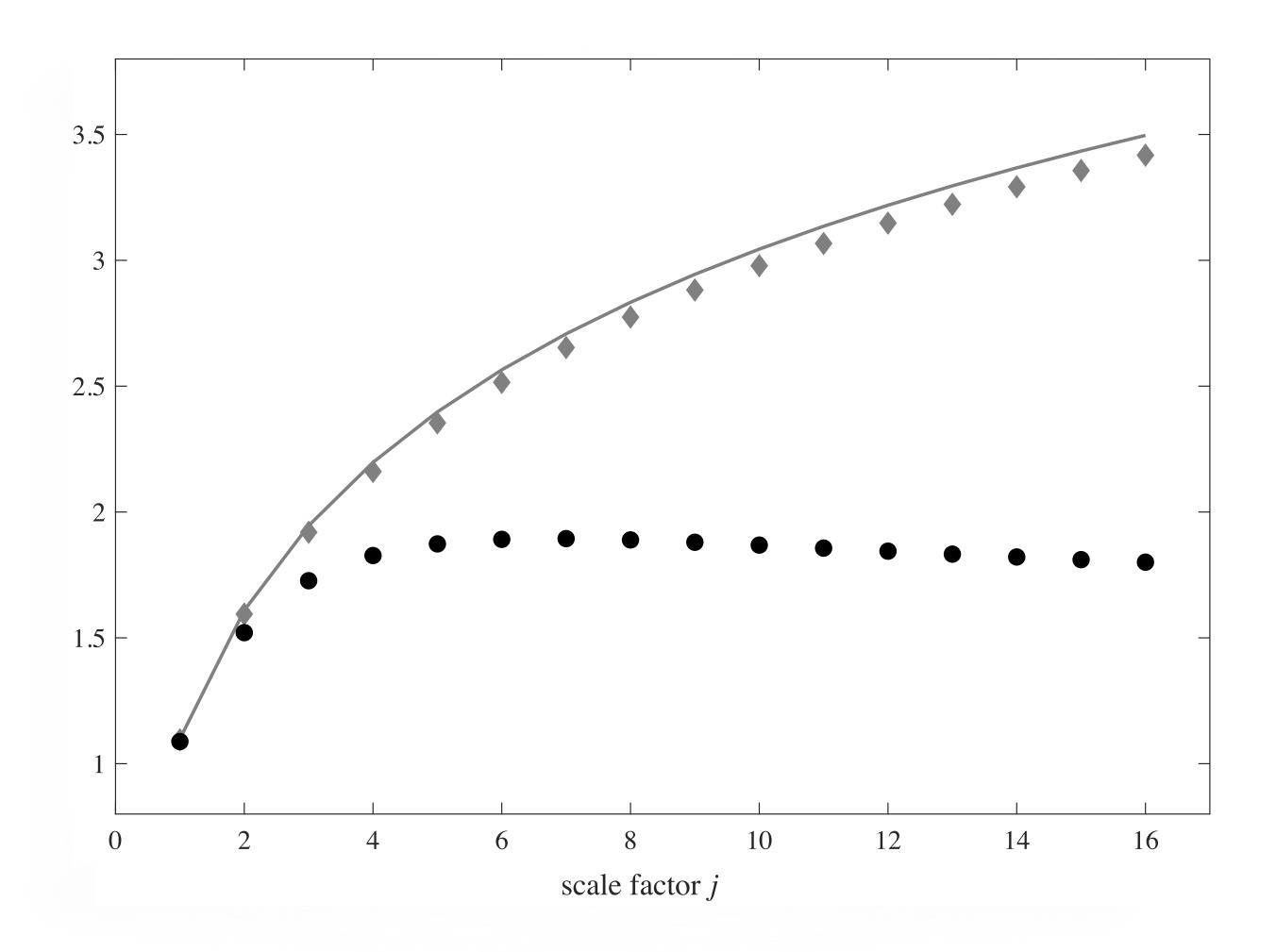


RECENT RESULTS

1. 3-sphere as emerging geometry

2. large fluctuations

3. large correlations





OPEN QUESTIONS IN QUANTUM GRAVITY

- What are the observables, when localization does not rely on background space and time?
- How should be think about time in this picture?
- What are the degrees of freedom?
- What's the interplay between the Planck scale and the cosmological scale?
- What is the role of quantum fluctuations of spacetime?

1. INTERPRETATION OF THE QUANTUM THEORY

■ Computing in quantum mechanics requires a PARTITION

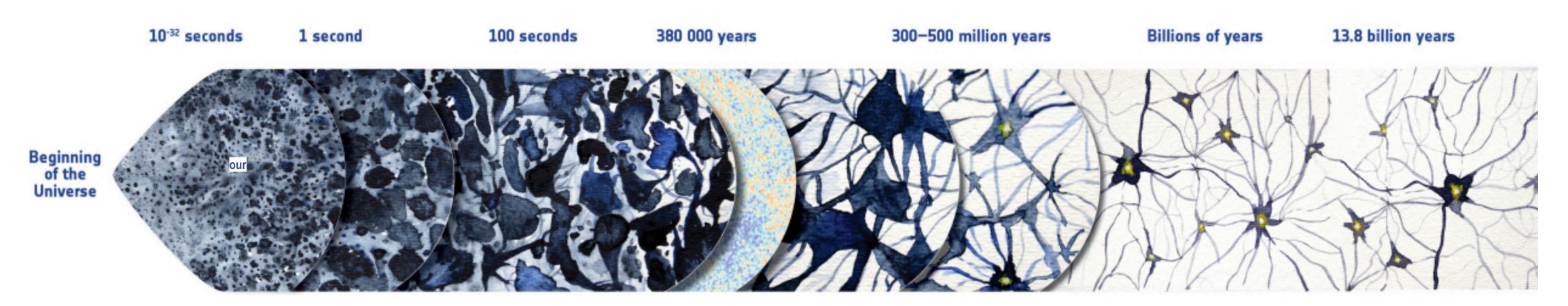


Image credit: ESA

2. WHERE IS SPACE?

- 1. "Space" in the most weak sense: the relation determined by the continuity of things
- 2. "Space" in the sense of geometry: quantity that we can measure with rods
- 3. "Space" in the sense of a continuous Riemaniann space: not in the theory
- 4. "Space" in the sense of the container within which things happen: not in the theory

3. WHERE IS TIME?

- 1. "Time" as change: variables change one with respect to another
- 2. "Time" as preferred time variable: not in the theory
- 3. "Time" in the sense of oriented flowing: compatible but not part of the theory

4. WHAT ARE THE OBSERVABLES?

■ In the considered application: area and volumes of space regions

■ LOCALIZATION: well defined relationally

■ NO need of continuous space

■ NO need of background space

5. WHAT ARE THE DEGREES OF FREEDOM?

■ The same as in General Relativity!

 \blacksquare quantum numbers j's and v's \longrightarrow $g_{ab}(x)$

■ CLASSICAL LIMIT in Quantum Gravity

analog to: photons — Maxwell equations

6. PLANCK VS COSMOLOGICAL SCALE

■ DISCRETIZATION ≠ QUANTUM DISCETENESS

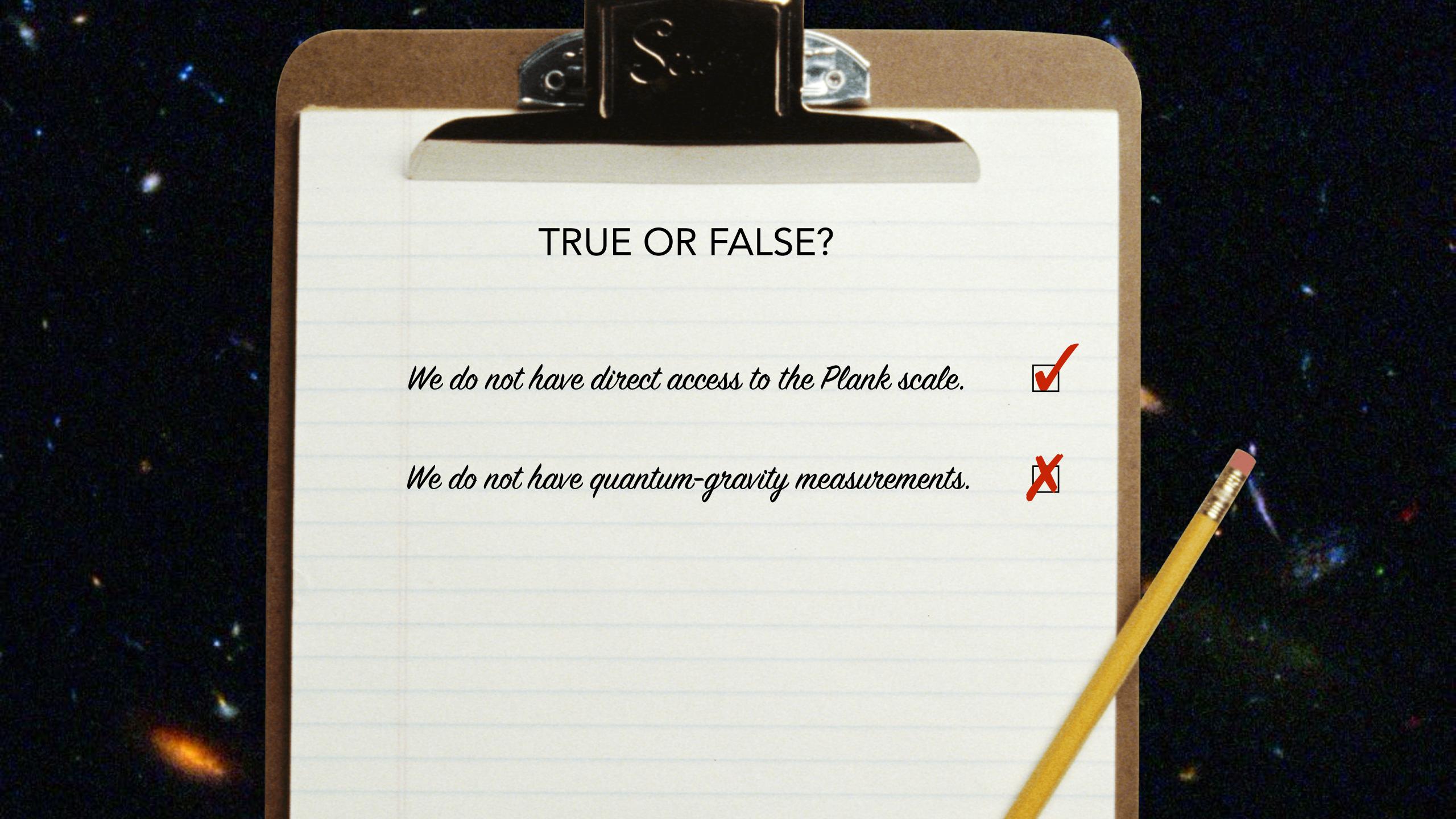
■ GRAPH TRUNCATION = APPROXIMANTION

■ QUANTUM COSMOLOGY: non-trivial interplay

7. WHAT IS THE ROLE OF QUANTUM FLUCTUATIONS?

■ SUPERPOSITION: generic states of the geometry are superposed ones





We do not have quantum-gravity measurements.



- Supersymmetric particles
 Violation Lorentz Invariance
 QG imprint on initial cosmological perturbations
- Cosmological variation of couplings
- Quantum decoherence and state collapse
- TeV Black Holes
- Generalized uncertainty principle
- Violation of discrete symmetries
- Speed of the gravitons
- Gravitational Wave Ech
- Planck scale spacetime fuzziness
- Entangled Masses?

